

# General Relativity

## Final Exam

4/11/2016

Please write your first and last name and your student number on the first page.

### Problem 1

Consider the two dimensional metric

$$ds^2 = \frac{1}{(1-r^2)^2} (dr^2 + r^2 d\theta^2)$$

with  $0 \leq r < 1$ .

1. Write the components of the metric and compute the Christoffel symbols.
2. How many independent components does the Riemann tensor have in 2 dimensions? Using these symmetries, compute the Ricci scalar of this manifold.
3. Derive the geodesic equations. Find the solution to the geodesic equation connecting a point  $r_0, \theta = 0$  to the point  $r_0, \theta = \pi$  and compute the length of this geodesic as a function of  $r_0$ .

### Problem 2

Consider the Schwarzschild metric

$$ds^2 = \left(1 - \frac{2GM}{r}\right) dt^2 - \left(1 - \frac{2GM}{r}\right)^{-1} dr^2 - r^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$

- i) Find the lowest bound for the radius a **stable** circular orbit for a massive particle.
- ii) Compute the period of the circular orbit at this lowest bound: a) according to an observer at infinity  
b) according to the proper time of an observer moving together with the orbiting particle.
- iii) Calculate the same for the innermost unstable circular orbit.

### Problem 3

In this problem we want to study how gravitational waves would behave in a 3-dimensional universe (as opposed to our 4-dimensional one). Consider the Einstein equations in 3 space-time dimensions, in the absence of matter.

- i) Expand the metric around 3d flat Minkowski space, and derive the linearized Einstein equations.
- ii) State what is the “gauge invariance” of the equations, and use it to simplify the linearized Einstein equations.
- iii) By following a similar analysis as the one in 4d, determine what are the possible polarizations of gravitational waves in 3d (Hint: the result has some surprising aspects).